Reading and Problem Assignments for Physics 243A Surface Physics of Materials: Spectroscopy, Fall, 2016 (In order of coverage in lecture)



Office hours:

TA: Galina Malovichko- 2:00-3:00 PM, Mondays, Physics 416

Instructor: Chuck Fadley, 2:30-3:30, Tuesdays and Thursdays, Physics 241 (others by special appointment)

Class consultant: Shih-Chieh Lin- 4:20-5:20, Wednesdays, Physics 221

Handing in problem sets:

Hand in problem sets in class, or in secure submission box with class number on it. The new homework submission boxes are located in the east wing at the first floor north entrance to the Physics Building – the same hallway where our student computer lab. is located (rm 106 Physics)

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				5.93 136	26 5.7	5.46 126	5.4 12	405 24.7	4.55 105	4.0 92	0	2.9 60	6				·							Ner I	e à		3	

death of the alaments

Cohesive energies and surface tension

ELEMENT	$\frac{E_{coh}}{\left(\frac{keal}{mal}\right)}$	AH _{EUS} (<u>kcal</u>)	Econ (90)
Si	107	9.47	8.9 %
Ca	42.1	2.23	5.3 % (P) (P)
Sc	90.6	3.80	4.2 70
Mn	68.7	3.45	5.0 00 000
Fe	98.9	3.56	3.6%
Cu	80.8	3.11	3.8%
Zn	34.1	1.60	5.190
Ge	89.3	8.30	9.396)

AHFUS < < ECOH BOND STRENGTHS/UR & BOND STRENGTHS/SOLID SOUD & O'LIQUID SOLID > YLIQUID



Fig. 1.4. Surface tension of the elements in the liquid phase (Schmit, 1974).

NOTE LOWER VALUES FOR HALF-FILLED HUND'S RULE STATES WITH REDUCED BONDING HYBRIDIBATION (d⁵, f⁷) AS WELL AS FOR FILLED-SHELL STATES



ENERGY TO PRODUCE EACH STEP

 $\gamma(\theta) = \gamma(0) + (\beta/a)|\theta|$ + BNBRGY OF INT. BETWEEN STEPS

The Wulff construction

Fig. 1.5. A vicinal surface.



Fig. 1.6. Polar plot of the surface tension at T = 0 (solid curve) and the Wulff construction of the equilibrium crystal shape (dashed curve) (Herring, 1951b).



7 Zangwill, p.13 Fig. 1.7. Electron micrograph of a lead crystal at 473 K (Heyraud & Metois, 1983).

Scanning electron microscopy Images of Pb crystallites



Wulff theorem: $\gamma_i/h_i = constant$

Fig. 1.8. Anisotropy of γ relative to $\langle 111 \rangle$ for lead as a function of temperature (Heyraud & Metois, 1983).



8 Zangwill, p.14

Scanning tunneling microscopy Images of Pb crystallites

a b (112)(100)(111) (221 C

C. Bombis et al., Surface Science 511 (2001) 83-96

Fig. 1. STM images of equilibrated Pb crystallite. (a) Vertical projection of complete crystallite, with the (111) facet parallel to the substrate, T = 323 K. Note (112) and (221) facets near the periphery of the (111) facet. Large (100) facets are visible at the contact line. Image size: $1300 \times 1300 \text{ nm}^2$. (b) Image of the (111) facet and vicinal range of same crystallite as in (a). A double tip effect is seen on the r.h.s. of the facet. Image size: $670 \times 670 \text{ nm}^2$. (c) 3D image of crystallite at T = 393 K, with the (100) facet rotated into the paper plane; center of the (100) facet approximately at the contact line. Due to the higher temperature, some bumps have been generated by local surface-to-tip contacts. Lateral dimension: 600 nm.

Scanning tunneling microscopy images of Pb crystallites

(a)





Fig. 2. (a) STM image of a section of the ECS of a Pb crystallite at 383 K. The image shows the facet to vicinal surface transition where the latter is characterized by a train of monatomic steps. (b) Example of a line scan across the ECS in the area of facet to vicinal surface transition. The line scan, $z(x) - z_0$ versus $x - x_0$, normalized relative to the facet edge at z_0 , x_0 , is fitted by Eq. (3) without the f_4 term.

Emundts et al., Surf. Sci. 481, 13 (2001)





Oxidation of Silicon: Time evolution at 580°C, Langmuir form



Langmuir-type adsorption: Growth rate is proportional to bare Si surface

$$\frac{d\theta}{dt} \propto P_{s}P(1-\theta)^{0}$$

 $\theta = \text{oxide coverage}$
 $P_{s} = \text{sticking probability}$
 $P = \text{oxygen pressure}$
 $n = \text{reaction order}$

20

The time evolution follows $n = 1 \rightarrow$ first-order Langmuirian

Sticking probability, P_S = 0.016—constant if no interaction between adsorbed species

Enta et al., PRB 57, 6294 (1998)

At Equilibrium!

Fig. 9.2. Adsorption isotherms for CO/Pd(111) (Ertl & Koch, 1970).



CO on Pd(111)- The 0.5 ML saturation structures-at equilibrium!



Coexistent- a slightly preferred at 120K



2 COs per rectangular cell, 4 COs per c(4x2) Fig. 5. c(4 × 2)-2CO models. (a) CO molecules occupying bridge sites. (b) CO molecules in fcc and hcp threefold hollow sites. A primitive cell of each c(4 × 2) structure is indicated.

Surface Science 512 (2002) 48-60



J. Electroanal. Chem., 353 (1993) 281–287 Pd(111)c(4×2)



Diffusion of water on TiO₂ and dissociation at an oxygen vacancy defect on the surface

Topics in Catalysis (2010) 53:423–430 doi 10.1007/s11244-010-9454-3



Other STM movies at: http://phys.au.dk/forskning/forskning somraader/condensed-matterphysics/spm/stm-movies/ **Fig. 4 a**–**b** Configuration and adsorption potential energy (PW91) of adsorbed monomeric water at Ti_{5f} sites close to an O_{br} vacancy. **c**–**d** Transition and final state of an adsorbed water molecule that jumps into an O_{br} vacancy. **e** Configuration with a pair of H adatoms in the O_{br} row after the water has dissociated. **f** Potential energy diagram. **g**–**j** STM images showing the diffusion of an adsorbed water molecule (*blue square*) that dissociates in an O_{br} vacancy (*white circle*) forming a pair of H adatoms (two neighboring *white dots*). STM images are extracted from movie "dissociation" published with [22]





X-ray spectrum from a rhodium target at 60 keV electron energy





X-Ray energies from the "X-Ray Data Booklet"

Element	Kα _l	Koz	Κβ 1	Laı	La ₂	Lβ ₁	Lβ ₂	L'n	Μα ₁
3 Li	54.3								
4 Be	108.5								
5 B	183.3								
6 C	277								
7 N	392.4		Popula	ar labor	atory so	urces			
8 O	524.9		/ for ph	otooloc	tronsna	octrosco	nv		
9 F	676.8	•			a on spe		, py		
10 Ne	848.6	848.6							
11 Na	1,040.98	1,040.98 🖌	1,071.1						
12 Mg	1,253.60	1,253.60	1,302.2						
13 Al	1,486.70	1,486.27	1,557.45						
14 Si	1 ,739.98	1,739.38	1,835.94						
15 P	2,013.7	2,012.7	2,139.1						
16 S	2,307.84	2,306.64	2,464.04						
17 Cl	2,622.39	2,620.78	2,815.6						
18 Ar	2,957.70	2,955.63	3,190.5						
19 K	3,313.8	3,311.1	3,589.6						
20 Ca	3,691.68	3,688.09	4,012.7	341.3	341.3	344.9			
21 Sc	4,090.6	4,086.1	4,460.5	395.4	395.4	399.6			

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Table 1-2. Photon energies, in electron volts, of principal K-, L-, and M-shell emission lines.

X-Ray energies from the "X-Ray Data Booklet" (cont'd.)

Element	Ka _l	Ka ₂	Κ β ₁	$L\alpha_1$	La ₂	Lβ ₁	Lβ ₂	Lŋ	Mα ₁
22 Ti	4,510.84	4,504.86	4,931.81	452.2	452.2	458.4			
23 V	4,952.20	4,944.64	5,427.29	511.3	511.3	519.2			
24 Cr	5,414.72	5,405.509	5,946.71	572.8	572.8	582.8			
25 Mn	5,898.75	5,887.65	6,490.45	637.4	637.4	648.8			
26 Fe	6,403.84	6,390.84	7,057.98	705.0	705.0	718.5			
27 Co	6,930.32	6,915.30	7,649.43	776.2	776.2	791.4			
28 Ni	7,478.15	7,460.89	8,264.66	851.5	851.5	868.8			
29 Cu	8,047.78	8,027.83	8,905.29	929.7	929.7	949.8			
30 Zn	8,638.86	8,615.78	9,572.0	1,011.7	1,011.7	1,034.7			
31 Ga	9,251.74	9,224.82	10,264.2	1,097.92	1,097.92	1,124.8			
32 Ge	9,886.42	9,855.32	10,982.1	1,188.00	1,188.00	1,218.5			
33 As	10,543.72	10,507.99	11,726.2	1,282.0	1,282.0	1,317.0			
34 Se	11,222.4	11,181.4	12,495.9	1,379.10	1,379.10	1,419.23			
35 Br	11,924.2	11,877.6	13,291.4	1,480.43	1,480.43	1,525.90			
36 Kr	12, 649	12, 59 8	14,112	1,586.0	1,586.0	1,636.6			
37 Rb	13,395.3	13,335.8	14,961.3	1,694.13	1,692.56	1,752.17			
38 Sr	14,165	14,097.9	15,835.7	1,806.56	1,804.74	1,871.72			
39 Y	14,958.4	14,882.9	16,737.8	1,922.56	1,920.47	1,995.84			
40 Zr	15,775.1	15,690.9	17,667.8	2,042.36	2,039.9	2,124.4	2,219.4	2,302.7	
41 Nb	16,615.1	16,521.0	18,622.5	2,165.89	2,163.0	2,257.4	2,367.0	2,461.8	
42 Mo	17,479.34	17,374.3	19,608.3	2 ,293 .16	2,289.85	2,394.81	2,518.3	2,623.5	111,323 1
43 Tc	18,367.1	18,250.8	20,619	2,424	2,420	2,538	2,674	2,792	19 - A
44 Ru	19,279.2	19,150.4	21,656.8	2,558.55	2,554.3 1	2,683.23	2,836.0	2,964.5	• •
45 Rh	20,216.1	20,073.7	22,723.6	2,696.74	2,692.05	2,834.41	3,001.3	3,143.8	
46 Pd	21,177.1	21,020.1	23,818.7	2,838.61	2,833.29	2,990.22	3,171.79	3,328.7	
47 Ag	22,162.92	21,990.3	24,942.4	2,984.31	2,978.21	3,150.94	3,347.81	3,519.59	
48 Cd	23,173.6	22,984.1	26,095.5	3,133.73	3,126.91	3,316.57	3,528.12	3,716.86	
49 In	24,209.7	24,002.0	27,275.9	3,286.94	3,279.29	3,487.21	3,713.81	3,920.81	
50 Sn	25,271.3	25,044.0	28,486.0	3,443.98	3,435.42	3,662.80	3,904.86	4,131.12	
51 Sb	26,359.1	26,110.8	29,725.6	3,604.72	3,595.32	3,843.57	4,100.78	4,347.79	
52 Te	27,472.3	27,201.7	30,995.7	3,769.33	3,758.8	4,029.58	4,301.7	4,570.9	
53 I	28,612.0	28,317.2	32,294.7	3,937.65	3,926.04	4,220.72	4,507.5	4,800.9	
54 Xe	29,779	29,458	33,624	4,109.9	_	_		_	
55 Cs	30,972.8	30,625.1	34,986.9	4,286.5	4,272.2	4,619.8	4,935.9	5,280.4	
56 Ba	32,193.6	31,817.1	36,378.2	4,466.26	4,450.90	4,827.53	5,156.5	5,531.1	
57 La	33,441.8	33,034.1	37,801.0	4,650.97	4,634.23	5,042.1	5,383.5	5,788.5	833
58 Ce	34,719.7	34,278.9	39,257.3	4,840.2	4,823.0	5,262.2	5,613.4	6,052	883
59 Pr	36,026.3	35,550.2	40,748.2	5,033.7	5,013.5	5,488.9	5,850	6,322.1	929
60 Nd	37,361.0	36,847.4	42,271.3	5,230.4	5,207.7	5,721.6	6,089.4	6,602.1	978
61 Pm	38,724.7	38,171.2	43,826	5,432.5	5,407.8	5,961	6,339	6,892	
62 Sm	40,118.1	39,522.4	45,413	5,636.1	5,609.0	6,205.1	6,586	7.178	1.081

4~

2











In the central radiation cone:

 $\frac{\Delta \omega}{\omega} \approx \frac{1}{N}$ $\theta_{cen} \simeq \frac{1}{\sqrt{N}}$

5th

6th

 λ_{x} as observed will increase if the magnetic field B_o is increased and/or the gap of the magnets is made smaller, or if viewed away from the axis by an angle θ Odd harmonics are much stronger than even harmonics

ALS



 ϵ_{c} = Critical energy [keV] = 0.665 E²[GeV]B[T]



In practical units [photons $s^{-1} mr^{-2} (0.1\% bandwidth)^{-1}],$

$$\frac{d^2 \check{S}_{\rm B}}{d \theta d \psi} \bigg|_{\psi=0} = 1.327 \times 10^{13} E^2 [\text{GeV}] I[\text{A}] H_2(\psi) \quad .$$

1 mrad = 10⁻²(360/2\pi) = 0.0572°

The function $H_2(y)$ is shown in Fig. 2-1.



Fig. 2-1. The functions $G_1(y)$ and $H_2(y)$, where y is the ratio of photon energy to critical photon energy.

The distribution integrated over ψ is given by

In practical units [photons \cdot s⁻¹·mr⁻¹·(0.1% bandwidth)⁻¹],

$$\frac{d S_{\rm B}}{d \theta} = 2.457 \times 10^{13} E [{\rm GeV}] I[{\rm A}] G_1(y) \quad . \label{eq:generalized}$$

A L S

Variable polarization with a bend magnet: above and below plane



Advanced Light Source--Sasaki-Carr **Elliptically-Polarized Undulator:** Variable light polarization LCP Horiz. Lin. **RCP** Vert. Lin.

[Can also vary polarization to LCP, RCP by going above and below the orbit plane in a bend magnet]



Translating magnet arrays

$\lambda_{x}(A) = 12,398/[hv(eV)]$

Vacuum Soft x- "Tender" Ultraviolet rays x-rays (VUV)- ~200-2000 ~2000-~8-200 eV eV 10000 eV

THE ELECTROMAGNETIC SPECTRUM



Advanced Light Source--Typical Soft X-Ray Spectroscopy Beamline Layout: to ca. 1500 eV



 $n\lambda = d(sin(i) - sin(r))$

Soleil (Paris)—Typical hard x-ray spectroscopy beamline



MULTI-TECHNIQUE SPECTROMETER/ DIFFRACTOMETER (MTSD)

> Scienta electron spectrometer (hidden)

ALS BL 9.3.1 hv = 2-5 keV



Sample prep. chamber: LEED, Knudsen cells, electromagnet,...

Scienta soft x-ray spectrometer

Permits using all relevant soft and hard x-ray spectroscopies on a single sample: PS, PD, PH; XAS (e⁻ or photon detection), XES/RIXS, with MCD, MLD



Synchrotron radiation sources of the world- about 41 and growing Free-electron laser (UV, X-ray)- about 5 and growing



Nature Photonics 9, 281 (2015)


ALS Beamlines January 2014

Superbend

Beamlines





The Next Generation: The Free-Electron Laser





Fig. 2. Average brightness comparisons of the LCLS and other light sources, including proposed FELs at Brookhaven [14] and DESY [15].

"X-Ray Data Booklet"₃₉ See Fig. 2.10



[4, modified].

higher than 3rd generation synchrotron facilities

"X-Ray Data Booklet"₄₀ Fig. 2.10 Future – Ultrafast X-ray Science X-ray Free-Electron Lasers!

SLAC National Accelerator Laboratory





19 Dispanse

s.

The vacuum ultraviolet, soft x-ray, hard x-ray measurements:

The spectroscopies: free

hν

hν

PHOTOELECTRON SPECTROSCOPY= PHOTOEMISSION – PS, PES, UPS, XPS + DIFFRACTION-XPD, PhD + HOLOGRAPHY-PH + MICROSCOPY-PEEM



X-RAY ABSORPTION SPECTROSCOPY- XAS *NEAR-EDGE – NEXAFS, XANES + X-RAY MAGNETIC CIRCULAR/LINEAR DICHROISM- XMCD, XMLD #EXTENDED- EXAFS, XAFS



X-RAY EMISSION (FLUORESCENCE) SPECTROSCOPY

+ AUGER ELECRON SPECTROSCOPY (Always accompanies photoelectron emission) 43





(E.G. See pp. 1-38, 1-44, 5-18-5-19 in X-Ray Data Booklet)

Index of refraction = $n = 1 - \delta - i\beta$ (Sometimes with + signs on δ and/or β) δ = + no. = refractive decrement << 1 (Sometimes changes sign through absorption resonances) β = + no. = absorptive decrement << 1

 δ and β linked by Kramers-Kronig transform

n also = 1 – $(r_e/2\pi)\lambda_{hv}^2 \sum n_i f_i$ (0=fwd. scatt.)

 r_{e} = classical electron radius $= e^{2/4} \pi \epsilon_0 m_e^2 = 2.817 \times 10^{-15} m_e^2$ $\lambda_{hv} = x$ -ray wavelength

 $n_i = n_i$ atoms per unit volume

 f_1 = x-ray scattering factor for ith type of atom, in forward direction

Exponential absorption length = $I_{abs} = \lambda_{hv}/(4\pi\beta) = \Lambda_{hv}$

 θ_{CRIT} = critical grazing angle at which reflectivity begins ($R \approx 0.20$) = [28]^{0.5}

Sections 1.6 and 1.7 of X-Ray Data Booklet

Online data and calculations at: http://henke.lbl.gov/optical_constants/getdb2.html



1.7 ATOMIC SCATTERING FACTORS

Eric M. Gullikson

The optical properties of materials in the photon energy range above about 30 eV can be described by the atomic scattering factors. The index of refraction of a material is related to the scattering factors of the individual atoms by

$$n = 1 - \delta - i\beta = 1 - \frac{r_{\rm e}}{2\pi} \lambda^2 \sum_i n_i f_i(0) \quad , \tag{1}$$

where r_e is the classical electron radius, λ is the wavelength, and n_i is the number of atoms of type *i* per unit volume. The parameters δ and β are called the refractive index decrement and the absorption index, respectively. The complex atomic scattering factor for the forward scattering direction is

$$f(0) = f_1 + if_2 \quad . \tag{2}$$

The imaginary part is derived from the atomic photoabsorption cross section:

$$f_2 = \frac{\sigma_a}{2r_e\lambda} \quad . \tag{3}$$

The real part of the atomic scattering factor is related to the imaginary part by the Kramers-Kronig dispersion relation:

$$f_1 = Z^* + \frac{1}{\pi r_e hc} \int_0^\infty \frac{\varepsilon^2 \sigma_{\rm a}(\varepsilon)}{E^2 - \varepsilon^2} d\varepsilon \qquad (4)$$

In the high-photon-energy limit, f_1 approaches Z^* , which differs from the atomic number Z by a small relativistic correction:

$$Z^* \approx Z - (Z/82.5)^{2.37} \quad . \tag{5}$$

Sections 1.6 and 1.7 of X-Ray Data Booklet Plus the "Bible" of Soft X-Ray Optics: Henke, Gullikson, Davis, Atomic and Nuclear Data Tables 54, 181-342 (1993) The mass absorption coefficient μ (cm²/g) is related to the transmitted intensity through a material of density ρ (g/cm³) and thickness d by

$$I = I_0 e^{-\mu\rho d} \qquad (1)$$

Thus, the linear absorption coefficient is μ_{ℓ} (cm⁻¹) = $\mu\rho$. For a pure material, the mass absorption coefficient is directly related to the total atomic absorption cross section σ_a (cm²/atom) by

$$\mu = \frac{N_{\rm A}}{A} \sigma_{\rm a} \quad , \quad 4\pi\beta/\lambda_{\rm X} = \mu_{\ell} = (N_{\rm A}/A)\rho\sigma_{\rm a} = 1/\Lambda_{\rm hv} \quad (2)$$

where N_A is Avogadro's number and A is the atomic weight. For a compound material, the mass absorption coefficient is obtained from the sum of the absorption cross sections of the constituent atoms by

$$\mu = \frac{N_{\rm A}}{MW} \sum_{i} x_i \sigma_{\rm ai} \quad , \tag{3}$$

where the molecular weight of a compound containing x_i atoms of type *i* is $MW = \sum_i x_i A_i$. This approximation, which neglects interactions among the atoms in the material, is generally applicable for photon energies above about 30 eV and sufficiently far from absorption edges.

ENHANCED SURFACE SENSITIVITY @ GRAZING

SOME X-RAY OPTICAL EFFECTS: REDUCED PENETRATION DEPTHS AND INCREASED REFLECITIVITY AT GRAZING INCIDENCE ANGLES

 $\label{eq:GRIT} \begin{array}{l} \textbf{θ_{CRIT} = Grazing angle at which} \\ \textbf{$reflectivity begins} \\ \textbf{$(R \approx 0.20)$} \end{array}$

= [2δ]^{0.5}

Calculated online from: http://henke.lbl.gov/optical _constants/atten2.html



Angle (deg)

Angle (deg)

Multilayer Reflectivity



http://henke.lbl.gov/optical_constants/multi2.html

Multilayer Reflectivity

 $\rm Si/Mo$ d=3.nm s=0.nm N=20 at 1000.eV, P=-1.



What properties do wave functions of overlapping (thus indistinguishable) particles have?—electrons as example:



 $\psi = \psi(\vec{r}_1, \vec{s}_1; \vec{r}_2, \vec{s}_2)$, including spin of both electrons But labels can't affect any measurable quantity. E.g. – probability density :

 $|\psi(\vec{r}_1, \vec{s}_1; \vec{r}_2, \vec{s}_2)|^2 = |\psi(\vec{r}_2, \vec{s}_2; \vec{r}_1, \vec{s}_1)|^2$

Therefore

 $\psi(\vec{r}_{1}, \vec{s}_{1}; \vec{r}_{2}, \vec{s}_{2}) = \pm 1 \psi(\vec{r}_{2}, \vec{s}_{2}; \vec{r}_{1}, \vec{s}_{1})$ $\equiv \widehat{P}_{12} \psi(\vec{r}_{1}, \vec{s}_{1}; \vec{r}_{2}, \vec{s}_{2})$

with \hat{P}_{12} = permutation operator $\rightarrow \vec{r}_1, \vec{s}_1; \vec{r}_2, \vec{s}_2$ and eigenvalues of ± 1

Finally, all particles in two classes :FERMIONS : (incl. e⁻'s): ψ antisymmetric $s = \frac{1}{2}, \frac{3}{2}, \frac{5}{2}...$ $\hat{P}_{12}\psi = -1\psi$ BOSONS : (incl. photons): ψ symmetrics = 0, 1, 2, ... $\hat{P}_{12}\psi = +1\psi$

A brief review of electronic structure in atoms, molecules, and solids

Probability of finding two electrons at the same point in space with the same spin is zero: "the Fermi Hole"



Antisymmetry and the Pauli Exclusion Principle:

Try Helium, 2 electrons in ground state 1s wave functions, "1s²"

Simple normalized antisymmetric trial wave function is

$$\begin{split} \psi(\vec{r}_{1},\vec{s}_{1};\vec{r}_{2},\vec{s}_{2}) &= \frac{1}{\sqrt{2}} \Big[\varphi_{1s}(\vec{r}_{1},\vec{s}_{1}=\uparrow)\varphi_{1s}(\vec{r}_{2},\vec{s}_{2}=\downarrow) - \varphi_{1s}(\vec{r}_{1},\vec{s}_{1}=\downarrow)\varphi_{1s}(\vec{r}_{2},\vec{s}_{2}=\uparrow) \Big] \\ int erchanging labels gives \\ \psi(\vec{r}_{2},\vec{s}_{2};\vec{r}_{1},\vec{s}_{1}) &= \frac{1}{\sqrt{2}} \Big[\varphi_{1s}(\vec{r}_{2},\vec{s}_{2}=\uparrow)\varphi_{1s}(\vec{r}_{1},\vec{s}_{1}=\downarrow) - \varphi_{1s}(\vec{r}_{2},\vec{s}_{2}=\downarrow)\varphi_{1s}(\vec{r}_{1},\vec{s}_{1}=\uparrow) \Big] \\ &= -\psi(\vec{r}_{1},\vec{s}_{1};\vec{r}_{2},\vec{s}_{2}), \text{ as required} \end{split}$$

Can't tell which electron is spin up--indistinguishable

Also, if we try to put both electrons in 1s with spin-up (\uparrow), first term always cancels second term, and $\psi = 0$? Therefore, we have the Pauli Exclusion Principle



Basic energetics—Many e⁻ picture



Basic energetics—Many e⁻ picture

Photoelectron emission: $n\ell j \rightarrow photoelectron at E_{kinetic}$

$$hv = E_{binding}^{Vacuum} + E_{kinetic} = E_{binding}^{Fermi} + \varphi_{spectrometer} + E_{kinetic}$$

 $E_{binding}^{Vacuum}(Qn\ell j,K) = E_{final}(N-1,Qn\ell j hole,K) - E_{initial}(N)$



Basic energetics—Many e⁻ picture

Photoelectron emission: $n\ell j \rightarrow photoelectron at E_{kinetic}$



Koopmans' Theorem Calculation of C 1s Chemical Shifts in Small C-Containing Molecules



Figure 18 -- Plot of carbon ls binding energies calculated via Koopmans' Theorem against experimental binding energies for several carbon-containing gaseous molecules. For some molecules, more than one calculated value is presented. The slope of the straight line is unity. The two scales are shifted with respect to one another by 15 eV, largely due to relaxation effects. All of the theoretical calculations were of roughly double-zeta accuracy or better. (From Shirley, reference 7.)

> "Basic Concepts of XPS" Figure 18

Correlation and screening effects beyond a single Slater determinant: configuration interaction:

For general N-electron state K: $\Psi_{\kappa}(N) =$ weighted sum of Slater determinants Φ_{j} $=\sum_{i} C_{j\kappa} \Phi_{j}$

with probability of each $|C_{jK}|^2$

For example, for Ne, a highly accurate CI calculation by Barr involving 1071 distinct configurations of spatial orbitals¹²³ yields the following absolute values for the coefficients multiplying the various members of a few more important configurations: $\Phi_1 = 1s^22s^22p^6 = \text{Hartree-Fock}$ configuration— 0.984; $\Phi_2 = 1s^22s^12p^63s^1$ —0.005; $\Phi_3 = 1s^22s^22p^53p$ —0.009; $\Phi_4 = 1s^22s^22p^44p^2$ —0.007–0.030; and $\Phi_5 = 1s^22s^22p^43p4p$ —0.007–0.022. Approximately 70 distinct configurations have coefficients larger than 0.010 in magnitude, but only that for Φ_1 is larger than 0.030.

In density-functional theory (DFT)→local-density approximation (LDA):

Effects of non-local exchange

$$K_{ij} \equiv \left\langle \phi_i(\vec{r}_1) \, | \, \hat{K}_j \, | \, \phi_i(\vec{r}_1) \right\rangle = \iint \phi_i^*(\vec{r}_1) \phi_j^*(\vec{r}_2) \frac{1}{r_{12}} \phi_j(\vec{r}_1) \phi_i(\vec{r}_2) dV_1 dV_2$$

and additional correlation effects are replaced in solving for ground-state one-electron orbitals by a <u>local</u> exchange-correlation potential:

$$V_{ec,\uparrow(\downarrow)}(\vec{r}) \approx -\frac{3}{4} \left[\frac{3\rho_{\uparrow(\downarrow)}(\vec{r})}{\pi} \right]^{1/3} - \frac{0.056\rho_{\uparrow(\downarrow)}^{1/3}(\vec{r})}{0.079 + \rho_{\uparrow(\downarrow)}^{1/3}(\vec{r})}$$

Where $\rho_{\uparrow(\downarrow)}(\vec{r}) = \sum_{i=1,\dots,N} \phi_{i,\uparrow(\downarrow)}^*(\vec{r})\phi_{i,\uparrow(\downarrow)}(\vec{r}) = \sum_{i=1,\dots,N} \left| \phi_{i,\uparrow(\downarrow)}(\vec{r}) \right|^2$

+ Various corrections/approximations going beyond this: generalized gradient approximation (GGA), GW and GW + cumulant expansions, dynamical mean field theory (DMFT), hybrid functionals mixing Hartree-Fock and LDA, quantum Monte Carlo,...



• USE SEPARATION OF VARIABLES :

 $\Psi(r, o, \phi) = \mathcal{R}(r) \Theta(o) \overline{\Phi}(\phi)$

-ASSUMED FORM-

· SUBSTITUTE, REARBANGE ----

$$\phi: \frac{d^{2}\overline{\Phi}}{d\phi^{2}} + C_{\phi}\overline{\Phi} = 0 \implies \overline{\Phi}(\phi) = A e^{\pm iC_{\phi}\phi} \qquad ()$$

$$\stackrel{W_{1}}{=} 0, \pm 4, \dots \qquad \overline{\Phi}(\phi) = \sqrt{2\pi} e^{\pm im_{E}\phi} - c_{OH}, \dots \\ \underset{M_{2}}{=} 0, \pm 4, \dots \qquad \underline{\Phi}(\phi) = \sqrt{2\pi} e^{\pm im_{E}\phi} - c_{OH}, \dots \\ \underset{M_{2}}{=} 0, \pm 4, \dots \qquad \underline{\Phi}(\phi) = \sqrt{2\pi} e^{\pm im_{E}\phi} - c_{OH}, \dots \\ \underset{M_{2}}{=} 0, \pm 4, \dots \qquad \underline{\Phi}(\phi) = \sqrt{2\pi} e^{\pm im_{E}\phi} = 0 \qquad (2)$$

$$e^{i(\phi)} = \sqrt{4} (im_{E}\phi) + \left[C_{\phi} - \frac{w_{1}^{2}}{\sin^{2}\phi}\right] \otimes = 0 \qquad (2)$$

$$e^{i(\phi)} = \sqrt{4} (im_{E}\phi) + \left[C_{\phi} - \frac{w_{1}^{2}}{\sin^{2}\phi}\right] \otimes = 0 \qquad (2)$$

$$e^{i(\phi)} = \sqrt{4} (im_{E}\phi) + \left[C_{\phi} - \frac{w_{1}^{2}}{\sin^{2}\phi}\right] \otimes = 0 \qquad (2)$$

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$$e^{i(\phi)} = \sqrt{4} (im_{E}\phi) + \left[C_{\phi} - \frac{w_{1}^{2}}{\sin^{2}\phi}\right] \otimes = 0 \qquad (2)$$

$$e^{i(\phi)} = \sqrt{4} (im_{E}\phi) + \left[C_{\phi} - \frac{w_{1}^{2}}{\sin^{2}\phi}\right] \otimes = 0 \qquad (2)$$

$$e^{i(\phi)} = \sqrt{4} (im_{E}\phi) + \left[C_{\phi} - \frac{w_{1}^{2}}{\sin^{2}\phi}\right] \otimes = 0 \qquad (2)$$

$$e^{i(\phi)} = \sqrt{4} (im_{E}\phi) + \left[C_{\phi} - \frac{w_{1}^{2}}{\sin^{2}\phi}\right] \otimes = 0 \qquad (2)$$

$$e^{i(\phi)} = \sqrt{4} (im_{E}\phi) + \left[C_{\phi} - \frac{w_{1}^{2}}{\sin^{2}\phi}\right] \otimes = 0 \qquad (2)$$

$$e^{i(\phi)} = \sqrt{4} (im_{E}\phi) + \left[C_{\phi} - \frac{w_{1}^{2}}{\sin^{2}\phi}\right] \otimes = 0 \qquad (2)$$

$$e^{i(\phi)} = \sqrt{4} (im_{E}\phi) + \left[C_{\phi} - \frac{w_{1}^{2}}{\sin^{2}\phi}\right] \otimes (im_{E}\phi) + \left[C_{\phi} -$$

59

Atomic orbitals:

32 =

COMPLEX, IF MZO TABLE 6.1 THE HYDROGEN ATOM FOR n = 1, 2, AND 3* (3 - 1 - 1, 2, AND 3*) NORMALIZED WAVE FUNCTION (0) 9 (4) Rng(r) nem (r,G, d)= 0 100 0 0 $\frac{2}{a_0^{3/2}}e^{-r/a_0}$ $\frac{1}{\sqrt{\pi} a_0^{3/2}} e^{-r/a_0}$ $\frac{1}{\sqrt{2\pi}}$ $\frac{1}{\sqrt{2}}$ 15 **2.5** 1 0 $\frac{1}{\sqrt{2\pi}}$ $\frac{1}{\sqrt{2}}$ **2.6** 1 $\frac{1}{\sqrt{2\pi}}$ $\frac{\sqrt{6}}{2}\cos\theta$ 1 ± 1 1 $\frac{1}{4\sqrt{2\pi}} \frac{1}{a_0^{3/2}} \left(2 - \frac{r}{a_0}\right) e^{-r/2a_0}$ $\frac{1}{2\sqrt{2} a_0^{3/2}} \left(2 - \frac{r}{a_0}\right) e^{-r/2a_0}$ $\frac{1}{2\sqrt{6} a_0^{3/2}} \frac{r}{a_0} e^{-r/2a_0}$ $\frac{1}{4\sqrt{2\pi}} \frac{r}{a_0^{3/2}} \frac{r}{a_0} e^{-r/2a_0} \cos \theta$ $\frac{\sqrt{3}}{2}\sin\theta$ $\frac{1}{\sqrt{2\pi}}e^{\pm i\phi}$ $\frac{1}{2\sqrt{6}a_0^{3/2}}\frac{r}{a_0}e^{-r/2a_0}$ $\frac{1}{8\sqrt{\pi}} \frac{r}{a^{3/2}} \frac{1}{a_0} e^{-r/2a_0} \sin \theta \ e^{\pm i\phi}$ 2831 $\begin{array}{ccc} 0 & \frac{1}{\sqrt{2\pi}} & \frac{1}{\sqrt{2}} \\ 0 & \frac{1}{\sqrt{2\pi}} & \frac{\sqrt{6}}{2}\cos\theta \end{array}$ $\frac{2}{81\sqrt{3} a_0^{3/2}} \left(27 - 18 \frac{r}{a_0} + 2 \frac{r^2}{a_0^2} \right) e^{-r/3a_0}$ $\frac{1}{81\sqrt{3\pi}} \frac{1}{a_0^{3/2}} \left(27 - 18 \frac{r}{a_0} + 2 \frac{r^2}{a_0^2} \right) e^{-r/3a_0}$ $\frac{4}{81\sqrt{6} a_0^{3/2}} \left(6 - \frac{r}{a_0} \right) \frac{r}{a_0} e^{-r/3a_0}$ $\frac{\sqrt{2}}{81\sqrt{\pi} a_0^{3/2}} \left(6 - \frac{r}{a_0} \right) \frac{r}{a_0} e^{-r/3a_0} \cos \theta$ 3Po $\frac{1}{\sqrt{2\pi}} \pm 1 \qquad \frac{1}{\sqrt{2\pi}} e^{\pm i\phi} \qquad \frac{\sqrt{3}}{2} \sin \theta$ $\frac{4}{81\sqrt{6}a_0^{3/2}}\left(6-\frac{r}{a_0}\right)\frac{r}{a_0}e^{-r/3a_0}$ $\frac{1}{81\sqrt{\pi}} \left(6 - \frac{r}{a_0} \right) \frac{r}{a_0} e^{-r/3a_0} \sin \theta \ e^{\pm i\phi}$ 3**P** 1 $0 \quad \frac{1}{\sqrt{2\pi}} \qquad \frac{\sqrt{10}}{4} (3 \cos^2 \theta - 1) \qquad \frac{4}{81\sqrt{30}} \frac{r^2}{a_0^{3/2}} \frac{r^2}{a_0^2} e^{-r/3a_0}$ $\frac{1}{81\sqrt{6\pi}} \frac{r^2}{a_0^{3/2}} e^{-r/3a_0} (3\cos^2\theta - 1)$

FOR

*The quantity $a_0 = 4\pi\epsilon_0 \hbar^2/me^2 = 5.3 \times 10^{-11}$ m is equal to the radius of the innermost Bohr orbit.

60 R

MANY

$\Psi_{n\ell m_\ell m_s}$	r,θ,φ,s	$(pin) = \Psi_{n\ell m_{\ell}}$	$r, \theta, \varphi) x [\alpha(\uparrow) \text{ or } \beta(\downarrow)]$	$\mathbf{Y}_{m}(\boldsymbol{\theta},\boldsymbol{\phi}) =$
The a	atom	ic orbit	als:	" spherical harmonics "
TABLE 6.1 NORMALIZED V		ONS OF THE HYDROGE		hem real for convenience
			$\frac{2}{r^{3/2}}e^{-r/a_0}$	$\frac{1}{\sqrt{1-\frac{1}{2}}e^{-r/a_0}} = \frac{1}{\sqrt{1-\frac{1}{2}}} = \frac{1}{\sqrt{1-\frac{1}{2}}}$
2 0 0 25 2 1 0	$\frac{1}{\sqrt{2\pi}}$	$\frac{1}{\sqrt{2}}$	$\frac{1}{2\sqrt{2}} \frac{1}{a_0^{3/2}} \left(2 - \frac{r}{a_0}\right) e^{-r/2a_0}$ $\frac{1}{1} = r - \frac{r}{2a_0}$	$\frac{\sqrt{\pi} a_0^{-r}}{4\sqrt{2\pi} a_0^{3/2}} \left(2 - \frac{r}{a_0}\right) e^{-r/2a_0}$
2 1 ±1 2P±1	$\frac{1}{\sqrt{2\pi}}e^{\pm i\phi}$	$\frac{\frac{1}{2}\cos\theta}{\frac{\sqrt{3}}{2}\sin\theta}$	$\frac{1}{2\sqrt{6} a_0^{3/2}} \frac{r}{a_0} e^{-r/2a_0}$ $\frac{1}{2\sqrt{6} a_0^{3/2}} \frac{r}{a_0} e^{-r/2a_0}$	$\frac{1}{8\sqrt{\pi} a_0^{3/2}} \frac{r}{a_0} e^{-r/2a_0} \sin \theta e^{\pm i\phi}$
3 1 0 3 2	$\frac{1}{\sqrt{2\pi}}$ $\frac{1}{\sqrt{2\pi}}$	$\frac{\frac{1}{\sqrt{2}}}{\frac{\sqrt{6}}{2}\cos\theta}$	$\frac{2}{81\sqrt{3}} \frac{r^2}{a_0^{3/2}} \left(27 - 18\frac{r}{a_0} + 2\frac{r^2}{a_0^2}\right) e^{-r/3a_0}$ $\frac{4}{81\sqrt{6}} \frac{r^2}{a_0^{3/2}} \left(6 - \frac{r}{a_0}\right) \frac{r}{a_0} e^{-r/3a_0}$	$\frac{1}{81\sqrt{3\pi} a_0^{3/2}} \left(27 - 18 \frac{r}{a_0} + 2 \frac{r^2}{a_0^2}\right) e^{-r/3a_0}$ $\frac{\sqrt{2}}{81\sqrt{\pi} a_0^{3/2}} \left(6 - \frac{r}{a_0}\right) \frac{r}{a_0} e^{-r/3a_0} \cos\theta \longrightarrow \text{ node for}$
$ \begin{array}{ccccccccccccccccccccccccccccccccc$	$\frac{1}{\sqrt{2\pi}} e^{\pm i\phi}$ $\frac{1}{\sqrt{2\pi}}$	$\frac{\sqrt{3}}{2}\sin\theta$ $\frac{\sqrt{10}}{(3\cos^2\theta-1)}$	$\frac{4}{81\sqrt{6} a_0^{3/2}} \left(6 - \frac{r}{a_0} \right) \frac{r}{a_0} e^{-r/3a_0}$ $\frac{4}{\sqrt{6} e^{-r/3a_0}} \frac{r^2}{2} e^{-r/3a_0}$	$\frac{1}{81\sqrt{\pi} a_0^{3/2}} \left(6 - \frac{r}{a_0} \right) \frac{r}{a_0} e^{-r/3a_0} \sin \theta \ e^{\pm i\phi} \qquad r = 6a_0$ $\frac{1}{81\sqrt{\pi} a_0^{3/2}} \left(6 - \frac{r}{a_0} \right) \frac{r}{a_0} e^{-r/3a_0} (3 \cos^2 \theta - 1) \rightarrow \text{nodes}$
3 2 ±1 34 ±1 3 2 ±2	$\frac{\sqrt{2\pi}}{\sqrt{2\pi}}e^{\pm i\phi}$	$\frac{4}{\sqrt{15}} \sin \theta \cos \theta$ $\sqrt{15}$	$81\sqrt{30} \ a_0^{3/2} \ a_0^2} \frac{4}{81\sqrt{30}} \ a_0^{3/2} \ a_0^2} \ e^{-r/3a_0} $	$81\sqrt{6\pi} a_0^{3/2} a_0^{2} \qquad \text{for } \cos^2\theta = \frac{1}{81\sqrt{\pi} a_0^{3/2}} \frac{r^2}{a_0^2} e^{-r/3a_0} \sin\theta\cos\theta e^{\pm i\phi} \qquad 1/3 \\ \theta = 54.7^\circ,$
3d=2	$\frac{1}{\sqrt{2\pi}}e^{\pm 2i\phi}$	$\frac{110}{4}\sin^2\theta$	$\frac{1}{81\sqrt{30}} \frac{1}{a_0^{3/2}} \frac{1}{a_0^2} e^{-r/3a_0}$	$\frac{1}{162\sqrt{\pi} a_0^{3/2}} \frac{1}{a_0^2} e^{-r/3a_0} \sin^2 \theta e^{\pm 2i\phi} \qquad 125.3^{\circ}$

CHANGES FOR

MANY C ATOMS

*The quantity $a_0 = 4\pi\epsilon_0 \hbar^2/me^2 = 5.3 \times 10^{-11}$ m is equal to the radius of the innermost Bohr orbit.

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 $e^{-r/na_0} \rightarrow e^{-Zr/na_0}$ for hydrogenic Z \rightarrow Z_{eff}(r) in many-e⁻ atoms

Overall, in a many-electron system:

•Anti-symmetry of total wave function implies: Pauli Exclusion Principle: No two electrons can have all the same quantum nos. n, ℓ, m_{ℓ}, m_s or, if spin-orbit split n, ℓ, j, m_i

•Electronic structure determined by filling n, ℓ (or n, ℓ, j) levels from lowest to highest energy ($E_{n\ell}$ or $E_{n\ell j}$ from radial Schroedinger Eqn. with Z_{eff})

•Partially filled subshells n, ℓ (or n, ℓ, j) have their lowest energy when a maximum no. of electrons have parallel spins = highest total spin angular momentum = \hat{S} (Hund's First Rule), and then they couple to yield highest total orbital angular momentum = \hat{L} (Hund's Second Rule) 62



Filling the Atomic Orbitals:





Intraatomic electron screening in many-electron atoms--a simple model

NEUTRAL

POINT CHARGE (~ C') + SPHERICAL SHELLS OF C' CHARGE (~ ORBITS) AROUND POINT-CHARGE NUCLEUS = ~ ATOM :

In many-electron atoms: For a given n, s feels nuclear charge more than p, more than d, more than f

Yields $Z_{\text{eff}}(r)$ in simple picture, and lifts degeneracy on ℓ in hydrogenic atom

[CHARGE IN REAL ATOM SMEARED OUT, BUT STILL ROUGHLY IN RADIAL SHELLS]



Intraatomic electron screening in many-electron atoms--a selfconsistent Q.M. calculation

Plus radial oneelectron functions: $P_{n\ell}(r) \equiv rR_{n\ell}(r)$

General useful rule: $(n-\ell)$ maxima in radial probability density $P_{n\ell}(r)^2 \equiv r^2$ $R_{n\ell}(r)^2$





FIGURE 7.13 The sequence of quantum states in an atom. Not to scale.



\$	5 ²	Perio	odic Ta	ble, wi	th the	Outer I in Thei	Electron ir Groun	Confi d Stat	guratic es	ons of	Neutr	al P	1	2	P ³	P ⁴	P ^S	1s ²
13	Be⁴	The 1 and i	The notation used to describe the electronic configuration of atoms and ions is discussed in all textbooks of introductory atomic physics. The letters s p d — signify electrons having orbital angular															
8	$2s^2$	The mome	e letters s, p, d, signify electrons having orbital angular mentum 0, 1, 2, in units h : the number to the left of the									$\frac{ar}{1e}$ 2s ²	2p 2s	$s^2 2p^2$	$2s^{2}2p^{3}$	2s ² 2p	$\frac{4}{2s^22p^5}$	$2s^22p^6$
a ¹¹	Mg ¹²	letter super	denot	es the to the r	princip ight de	al qua notes t	ntum nu he num	umber ber of	of one electro	orbit, ns in t	, and tl the orb	ne Al' it.	³ Si	i ¹⁴	P ¹⁵	S ¹⁶	CI17	Ar ¹⁸
8	3s²	41	d²	d3	d"	d ⁵	d6	ď	d ⁸	d	ď	382	3p 3s	² 3p ²	3s²3p	³ 3s ² 3p	¹ 3s ² 3p ⁵	3s²3p ⁶
19	Ca ²⁰	Sc ²¹	Ti ²²	V ²³	Cr ²⁴	Mn ²⁵	Fe ²⁶	Co ²⁷	Ni ²⁸	Cu	²⁹ Zn	³⁰ Ga	³¹ G	e ³²	As ³³	Se ³⁴	Br ³⁵	Kr ³⁶
8	4s ²	$\frac{3d}{4s^2}$	$\frac{3d^2}{4s^2}$	$\frac{3d^3}{4s^2}$	3d ⁵ 4s	$\frac{3d^{5}}{4s^{2}}$	$\frac{3d^{6}}{4s^{2}}$	3d 7 4s ²	$\frac{3d^8}{4s^2}$	3 <i>d</i> 4s	10 3d 4s	¹⁰ ² 4s ²	4p 48	s ² 4p ²	4s ² 4p	³ 4s ² 4p	⁴ 4s ² 4p ⁵	4s ² 4p ⁶
b ³⁷	`Sr ³⁸	Y ³⁹	Zr ⁴⁰	Nb41	Mo ⁺	TC ⁴³	Ru ⁴⁴	Rh ⁴⁵	Pd ⁴⁶	Ag	47 Co	l ⁴⁸ In	¹⁹ S	n ⁵⁰	Sb ⁵¹	Te ⁵²	153	Xe ⁵⁴
8	5s ²	$\frac{4d}{5s^2}$	$4d^2$ $5s^2$	4d⁴ 5s	4d⁵ 58	4d ⁶ 58	$\frac{4d^{7}}{5s}$	4d* 5s	4d 10 -	4d 5s	10 4d 5s	¹⁰ ² 5s ²	5p 5	$s^2 5 p^2$	5s²5p	³ 5s ² 5p	⁴ 5s ² 5p ⁵	5s ² 5p ⁶
S ⁵⁵	Ba ⁵⁶	La ⁵⁷	Hf ⁷²	Ta ⁷³	W ⁷⁴	Re ⁷⁵	Os ⁷⁶	lr ⁷⁷	Pt ⁷⁸	Au	⁷⁹ Hç	3 ⁸⁰ TI ⁸	¹ P	b ⁸²	Bi ⁸³	Po ⁸⁴	At ⁸⁵	Rn ⁸⁶
8	6s ²	$5d$ $6s^2$	$4f^{14}$ $5d^2$ $6s^2$	$5d^3$ $6s^2$	$5d^4$ $6s^2$	$5d^{5}$ $6s^{2}$	$5d^6$ $6s^2$	5d ° -	5d ⁹ 6s	5d 6s	¹⁰ 5 <i>a</i> 6s	2 ¹⁰ 26s ²	6p 6	$s^2 6p^2$	6s ² 6p	³ 6s ² 6p	4 6s ² 6p ⁵	6s ² 6p ⁶
r ⁸⁷	Ra ⁸⁸	Ac ⁸⁹		4	41		•••	44	(50) ;	14	ING		•		٩	614	
	—)	6d	C 4j	e ⁵⁸ P ^{C2} 4j	r ⁵⁹ N C ³ 4	$\begin{array}{c c} \mathbf{Id}^{60} & \mathbf{F} \\ f^4 & 4 \end{array}$	rm^{61} Since f^5 Appendix f^5	$ \begin{array}{c c} \mathbf{m}^{62} & \mathbf{E} \\ f^6 & 4 \end{array} $	f^7	Gd ⁶⁴ 4 <i>f</i> ⁷	Tb ⁶⁵ 4 <i>f</i> ⁸	Dy ⁶⁶ 4f ¹⁰	Ho ⁶⁷ 4f ¹¹	Er 4 <i>f</i>	⁶⁸ T	m^{69} Y f^{13} 4	f^{14} L f h h h h h h h h h h	u^{71} f^{14}
8	18*	18"	6	² 6	s ² 6	s ² 6	5s ² 6s	s ² 6	8 ²	5 <i>a</i> 6s ²	$6s^2$	6s ²	6s ²	6 s	2 6	s ² 6	s^2 6	s ²
= E	RCE I	TION	5 T	h ⁹⁰ P 5	a^{91} L f^2 5	f^{3}	Np ⁹³ P 5f ⁵ 5	u^{94} A f^{6} 5	4m⁹⁵	Cm ⁹⁶ 5 <i>f</i> ⁷	Bk ⁹⁷	Cf ⁹⁸	Es ⁹⁹	Fr	n ¹⁰⁰ N	ld ¹⁰¹ N	lo ¹⁰² L	.r ¹⁰³

$$SPM'-ORBIT SPLITTING OF LEVELS:$$

$$STITE STREET SPLITTING OF LEVELS:$$

$$SFFECTIVE \overrightarrow{B} (NUCLEUS ANOUND e^{-}) \ll \overrightarrow{L}$$

$$H_{3-0} = \widehat{S}(r) \overrightarrow{L} \cdot \overrightarrow{S}$$

$$SPLITS ALL NL NL LEVELS NL_{j} = L + 4_{2} - 2L + 4_{2}$$

$$SPLITS ALL NL LEVELS NL_{j} = L - 4_{2} - 2L$$

$$MIXES SPIN + ORBITAL ANGULAR MOM.:$$

$$W_{nLjmj} = (-1)^{4} M_{3}m_{j} - 4_{2}(0) + (-2)^{4} M_{3}m_{j} + 4_{2}(0)$$

$$m_{s} = + \frac{1}{2} M_{s} M_{s} - \frac{1}{2} M_{s}$$

$$MITH C1 AND C2 TABULATED CLEBSCH-GORDAN OR WIGNER 3j SYMBOLS$$

$$70$$

	Some st	IN-ORBIT S	PLITTINGS: (IN eV)
2p ⁶ -	L= 13 (AQ)	28 (Ni)	46 (7d)
201 203/2	0.4	/7.8	157.0
34 10 - 3	L = 30 (Zn)	48 (Cd.)	64 (GL)
34 34 34	0.1	6.7	32.3
4514-3	= 74 (W)	BY (Pb)	92 (U)
45 45 8	2.2	7.0	64

INCREASE WITH Z FOR A GIVEN LEVEL. IN GENERAL:



X-Ray Data Booklet--Section 1.1 ELECTRON BINDING ENERGIES

The energies are given in eV relative to the <u>vacuum level</u> for the rare gases and for H_2 , N_2 , O_2 , F_2 , and CI_2 ; relative to the <u>Fermi level</u> for the metals; and relative to the <u>top of the valence bands</u> for semiconductors (and insulators).

Electronic configuration	Element	K 1s	L ₁ 2s	$L_2 2p_{1/2}$	L ₃ 2p _{3/2}	M ₁ 3s	M ₂ 3p _{1/2}	M3 3p3/2	_
ls	1 H	13.6							-
1 <i>s</i> ²	2 He	24.6*						Micc	inc
1s ² 2s	3 Li	54.7*							,
1s ² 2s ²	4 Be	111.5*					\mathbf{N}	vale	nce
1s ² 2s ² 2p	5 B	188*						DE	•
$1s^2 2s^2 2p^2$	6 C	284.2*	Va	lence le	evels		7	D.E.	5
$1s^2 2s^2 2p^3$	7 N	409.9*	37.3*	~ 9	~ 9		ornolato	$\boldsymbol{\lambda}$	
$1s^2 2s^2 2p^4$	8 O	543.1*	41.6*	~ 13	~ 13		erpolate	u,	
$1s^2 2s^2 2p^5$	9 F	696.7*	~ 45	~ 17	~ 17	ext	rapolate		
1s ² 2s ² 2p ⁶	10 Ne	870.2*	48.5*	21.7*	21.6*				
[Ne] 3s	11 Na	1070.8†	63.5†	30.65	30.81				
[Ne] 3s ²	12 Mg	1303.0†	88.7	49.78	49.50				
[Ne] 3s ² 3p	13 Al	1559.6	117.8	72.95	72.55				
[Ne] $3s^2 3p^2$	14 Si	1839	149.7*b	99.82	99.42				
[Ne] $3s^2 3p^3$	15 P	2145.5	189*	136*	135*	Val			
[Ne] $3s^2 3p^4$	16 S	2472	230.9	163.6*	162.5*				
[Ne] $3s^2 3p^5$	17 Cl	2822.4	270*	202*	200*				
[Ne] $3s^2 3p^6$	18 Ar	3205.9*	326.3*	250.6†	248.4*	29.3*	15.9*	15.7*	
[Ar] 4s	19 K	3608.4*	378.6*	297.3*	294.6*	34.8*	18.3*	18.3*	
[Ar] 4s ²	20 Ca	4038.5*	438.4†	349.7†	346.2*	44.3 †	25.4*	25.4†	
	21 Sc	4492	498.0*	403.6*	398.7*	51.1*	28.3*	28.3*	
	22 Ti	4966	560.9†	460.2*	453.8*	58.7*	32.6*	32.6*	
X-Ray Data Booklet--Section 1.1 ELECTRON BINDING ENERGIES

Element	K 1s	L ₁ 2s	$L_2 2p_{1/2}$	L ₃ 2p _{3/2}	M ₁ 3s	M ₂ 3p _{1/2}	M3 3p3/2	M4 3d3/2	M5 3d5/2	N ₁ 4s	$N_2 \; 4p_{1/2}$	N3 4p3/2	
23 V	5465	626.7†	519.8†	512.1†	66.3†	37.2*	37.2†						
24 Cr	5989	696.0†	583.8†	574.1*	74.1*	42.2*	42.2*						
25 Mn	6539	769.1†	649.9†	638.7†	82.3*	47.2†	47.2*						
26 Fe	7112	844.6†	719.9†	706.8†	91.3†	52.7†	52.7†	vale	nce le	veis			
27 Co	7709	925.1†	793.2†	778.1†	101.0†	58.9†	59.9†						
28 Ni	8333	1008.6†	870.0†	852.7†	110.8†	68.0†	66.2†						
29 Cu	8979	1096.7†	952.3†	932.7	122.5†	77.3†	75.1†						
30 Zn	9659	1196.2*	1044.9*	1021.8*	139.8*	91.4*	88.6*	10.2*	10.1*				
31 Ga	10367	1299.0*b	1143.2†	1116.4†	159.5†	103.5†	100.0†	18.7†	18.7†				
32 Ge	11103	1414.6*b	1248.1*b	1217.0*b	180.1*	124.9*	120.8*	29.8	29.2	Valo	nco lo		
33 As	11867	1527.0*b	1359.1*b	1323.6*b	204.7*	146.2*	141.2*	41.7*	41.7*				
34 Se	12658	1652.0*b	1474.3*b	1433.9*b	229.6*	166.5*	160.7*	55.5*	54.6*				
35 Br	13474	1782*	1596*	1550*	257*	189*	182*	70*	69*				
36 Kr	14326	1921	1730.9*	1678.4*	292.8*	222.2*	214.4	95.0*	93.8*	27.5*	14.1*	14.1*	
37 Rb	15200	2065	1864	1804	326.7*	248.7*	239.1*	113.0*	112*	30.5*	16.3*	15.3 *	
38 Sr	16105	2216	2007	1940	358.7†	280.3†	270.0†	136.0†	134.2*	38.9†	21.3	20.1*	
39 Y	17038	2373	2156	2080	392.0*b	310.6*	298.8*	157.7*	155.8†	43.8*	24.4*	23.1*	
40 Zr	17998	2532	2307	2223	430.3†	343.5†	329.8†	181.1*	178.8†	50.6†	28.5*	27.1*	
41 Nb	18986	2698	2465	2371	466.6†	376.1†	360.6†	205.0†	202.3†	56.4†	32.6†	30.8†	
42 Mo	20000	2866	2625	2520	506.3†	411.6†	394.0†	231.1*	227.9*	63.2*	37.6†	35.5†	
43 Tc	21044	3043	2793	2677	544*	447.6	417.7	257.6	253.9*	69.5*	42.3*	39.9*	
44 Ru	22117	3224	2967	2838	586.1*	483.5†	461.4†	284.2*	280.0†	75.0†	46.3†	43.2†	
45 Rh	23220	3412	3146	3004	628.1†	521.3*	496.5†	311.9*	307.2*	81.4*b	50.5†	47.3*	
46 Pd	24350	3604	3330	3173	671.6†	559.94	532.3*	340.5*	335.2*	87.1*b	55.7†a	50.9†	
47 Ag	25514	3806	3524	3351	719.0†	603.8†	573.0†	374.0†	368.3	97.0†	63.7*	58.3†	



PES = photoemission = photoelectron spectroscopy XAS = x-ray absorption spectroscopy AES = Auger electron spectroscopy XES = x-ray emission spectroscopy RIXS = resonant inelastic x-ray scattering / x-ray Raman scatt.

The Soft and Hard X-Ray Spectroscopies



PES = photoemission = photoelectron spectroscopy XAS = x-ray absorption spectroscopy AES = Auger electron spectroscopy XES = x-ray emission spectroscopy REXS/RIXS = resonant elastic/inelastic x-ray scattering⁷⁵ MATRIX ELEMENTS IN The Soft and Hard X-Ray Spectroscopies: DIPOLE LIMIT

• <u>Photoelectron spectroscopy/photoemission:</u> $\varphi_{f}(\text{free})$ hv $\varphi_{f}(\text{free})$ $I \propto \left| \hat{e} \cdot \langle \varphi_{f}(1) | \vec{r} | \varphi_{i}(1) \rangle \right|^{2}$ $\varphi_{i}(\text{bound})$

Core and valence photoemission

 $La_{0.6}Sr_{0.4}MnO_3$



The Soft and Hard X-Ray Spectroscopies



PES = photoemission = photoelectron spectroscopy XAS = x-ray absorption spectroscopy AES = Auger electron spectroscopy XES = x-ray emission spectroscopy REXS/RIXS = resonant elastic/inelastic x-ray scattering⁷⁸ MATRIX ELEMENTS IN The Soft and Hard X-Ray Spectroscopies: DIPOLE LIMIT



Variation of Near-Edge X-Ray Absorption Fine Structure (NEXAFS) with Atomic No. for Some 3d Transition Metals

J. Stohr, "NEXAFS Spectroscopy" (Spring, 1992), Stohr and Siegmann, "Magnetism: From **Fundamentals to Nanoscale Dynamics**" (Springer **Series in Solid-State** Sciences, 2006), **Chapter 9** Download from 243A website: http://physics.ucdavi s.edu/Classes/Physic s243A/XMCD.stohr.si egmann.abridged-foremailing.pdf



Variation of Near-Edge X-Ray Absorption Fine Structure (NEXAFS) for Different Polymers



SCANNING TRANSMISSION X-RAY MICROSCOPY OF POLYMER BLEND





A FINAL (RELATIVISTIC) ATOMIC INTERACTION: SPIN-ORBIT COUPLING

(a) An electron circles an atomic nucleus, as viewed from the frame of reference of the nucleus. (b) From the electron's frame of reference, the nucleus is circling it. The magnetic field the electron experiences as a result is directed upward from the plane of the paper. Th interaction between the electron's spin magnetic moment and this magnetic field leads to th phenomenon of spin-orbit coupling.





Magnetic Circular Dichroism in X-Ray Absorption (XMCD): Only happens because of the spin-orbit effect

$$Set H_{2} \xrightarrow{SPM-ORDIT SPLITTING OF LEVELS}:$$

$$Set H_{2} \xrightarrow{S} EFFECTIVE \overrightarrow{B} (NUCLEUS ANDUND e^{-}) \ll \overrightarrow{L}$$

$$\widehat{H}_{3,0} = \widehat{S}(r) \overrightarrow{L} \cdot \overrightarrow{S}$$

$$\circ SPLITS ALL NL LEVELS \xrightarrow{NL_{1}} L = L + V_{2} - 2L + 2L$$

$$\circ SPLITS ALL NL LEVELS \xrightarrow{NL_{2}} L = -V_{2} - 2L$$

$$\bullet MIXES SPIN + ORBITAL ANGULAR MOM.:$$

$$\widehat{V}_{nL_{2}m_{1}} = C_{1} \underbrace{V}_{nL_{2}m_{2}} - \underbrace{V}_{nL_{2}m_{2}} \underbrace{V}_{nL_{2}m_{2}} + \underbrace{V}_{nL_{2}m_{2}} \underbrace$$

WITH C1 AND C2 TABULATED CLEBSCH-GORDAN OR WIGNER 3j SYMBOLS

Example: <u>Photoelectron</u> spin polarization from spin-orbit coupling and circularly-polarized radiation—The Fano Effect



Photoelectron spin polarization from spin-orbit coupling and circularly-polarized radiation—The Fano Effect



Fano effect and spin polarization (SP) in core photoelectron spectra—expt.



Magnetic Circular Dichroism in X-Ray Absorption (XMCD)

J. Stohr, Journal of Magnetism and Magnetic Materials 200 (1999) 470-497







REXS/RIXS = resonant elastic/inelastic x-ray scattering^{*}

MATRIX ELEMENTS IN The Soft and Hard X-Ray Spectroscopies: DIPOLE LIMIT



a complex operator, no simple selection rules





O1s

"Basic Concepts of XPS" 94 Figure 1



Figure 2. Scheme of the Auger process. A valence-level involved Auger emission is illustrated here, but the two electrons involved also could have come from core level, ϵ_4 , provided $\epsilon_5 - 2\epsilon_4 > 0$.



Plus see pp. 92-93 in "Basic Concepts of XPS"



X-Ray Data Booklet--Section 1.1 ELECTRON BINDING ENERGIES

The energies are given in eV relative to the <u>vacuum level</u> for the rare gases and for H_2 , N_2 , O_2 , F_2 , and Cl_2 ; relative to the <u>Fermi level</u> for the metals; and relative to the <u>top of the valence bands</u> for semiconductors (and insulators).

	El	ectr figu	onic ment ration	ł	K 1s	L ₁	2s	L_2	2p _{1/2}	L	3 2p _{3/2}		M ₁ 3s	M	2 3p _{1/2}	M3 31	P3∕2	
1 <i>s</i>	com	1	Н		13.6							1						
1 <i>s</i> ²		2	He		24.6*											Mi	SS	inc
1s ² 2s		3	Li		54.7*													
1s ² 2s ²		4	Be		111.5*											va	ler	ICe
1s ² 2s ² 2	P	5	В		188*	Г									$\mathbf{\mathcal{L}}$			
1s ² 2s ² 2	p^2	6	С		284.2*		va	Ien	ce I	ev	eis				7		L.3	
1s ² 2s ² 2	p ³	7	Ν		409.9*	31	7.3*	,	~ 9		~ 9] Ini	torn	olato	$\boldsymbol{\lambda}$		
1s ² 2s ² 2	.p ⁴	8	0		543.1*	4	1.6*	~	13		~ 13			tra				
1s ² 2s ² 2	p5	9	F		696.7*	~ 4	5	~ (17		~ 17			ua	Julate	u	\mathbf{N}	
1s ² 2s ² 2	p ⁶	10	Ne		870.2*	48	8.5*	2	21.7*		21.6*							
[Ne] 3s		11	Na	1	070.8†	63	3.5†	2	30.65		30.81							
[Ne] 3s ²		12	Mg	1	303.0†	88	8.7	4	19.78		49.50							
[Ne] 3s ²	3р	13	Al	1	559.6	117	7.8	Ĵ	2.95		72.55	Г						
[Ne] 3s ² :	3p ²	14	Si	1	839	149	9.7*b	9	9.82		99.42							
[Ne] 3s ²	3p ³	15	Р	2	2145.5	189)*	13	36*		135*		Va	len	ce le	vels		
[Ne] 3s ²	3p ⁴	16	S	- 2	2472	230	0.9	16	53.6*		162.5*							
[Ne] 3s ²	3p ⁵	17	Cl	2	2822.4	270)*	20)2*		200*							
[Ne] 3s ² :	3p ⁶	18	Ar	3	205.9*	320	5.3*	25	50.6†		248.4*		29.3*		15.9*	15	.7*	
[Ar]	45	19	К	3	608.4*	378	8.6*	- 29	97.3*		294.6*		34.8*		18.3*	18	.3*	
[Ar]	4 <i>s</i> ²	20	Ca	4	038.5*	438	8.4†	34	1 9.7†		346.2†		44.3 🕴		25.4†	25	.4†	
		21	Sc	4	492	498	8.0*	40)3.6*		398.7*		51.1*		28.3*	28	.3*	
		22	Ti	4	966	560	0.9†	46	50.2†		453.8†		58.7†		32.6†	32	.698	



REXS/RIXS = resonant elastic/inelastic x-ray scattering

THE AUGER PROCESS



Figure 2. Scheme of the Auger process. A valence-level involved Auger emission is illustrated here, but the two electrons involved also could have come from core level, ϵ_4 , provided $\epsilon_5 - 2\epsilon_4 > 0$.

Z

SIMP



1-14

					5 01101 5101	,			
Element	K 1s	L ₁ 2s	$L_2 2p_{1/2}$	L ₃ 2p _{3/2}	M ₁ 3s	M ₂ 3p _{1/2}	M ₃ 3p _{3/2}	M ₄ 3d _{3/2}	$M_5 3d_{5/2}$
23 V	5465	626.7†	519.8†	512.1†	66.3†	37.2*	37.2†		
24 Cr	5989	696.0†	583.8†	574.1*	74.1†	42.2*	42.2†	Diff	4.0
25 Mn	6539	769.1†	649.9†	638.7†	82.3†	47.2†	47.2†	* DITI. = '	11.2
26 Fc	7112	844.6†	719.9†	706.8†	91.3†	52.7*	52.7†		
27 Co	7709	925.1†	793.2†	778.1†	101.0†	58.9†	59.9†		
28 Ni	8333	1008.6†	870.0†	852.7†	110.8†	68.0†	66.2†		
29 Cu	8979	1096.7†	952.3†	932.7	122.5†	77.3†	75.1†		
30 Zn	9659	1196.2*	1044.9*	1021.8*	139.8*	91.4*	88.6*	10.2^{*}	10.1*

Electron binding energies

Table 1-2. Energies of x-ray emission lines (continued).

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. . .

Element	Ka ₁	Κα2	Κβ 1	La ₁	Laz	L _{β1}	Lβ ₂	L'n	Mα ₁
22 Ti	4,510.84	4,504.86	4,931.81	452.2	452.2	458.4			
23 V	4,952.20	4,944.64	5,427.29	511.3	511.3	519.2			
24 Cr	5,414.72	5,405.509	5, 946 .71	572.8	572.8	582.8			
25 Mn	5,898.75	5,887.65	6,490.45	637.4	637.4	648.8	→ Dif	f. = 11.4	
26 Fe	6,403.84	6,390.84	7,057.98	705.0	705.0	718.5			
27 Co	6,930.32	6,915.30	7,649.43	776.2	776.2	791.4			
28 Ni	7,478.15	7,460.89	8,264.66	851.5	851.5	868.8			
29 Cu	8,047.78	8,027.83	8,905.29	929.7	929.7	949.8	Se	e Table	s 1.1
30 Zn	8,638.86	8,615.78	9,572.0	1,011.7	1.011.7	1.034.7	an	d 1.2 in	X-Ray

Data Booklet

MATRIX ELEMENTS IN The Soft and Hard X-Ray Spectroscopies: DIPOLE LIMIT



1.3 FLUORESCENCE YIELDS FOR K AND L SHELLS

Jeffrey B. Kortright

Fluorescence yields for the K and L shells for the elements $5 \le Z \le 110$ are plotted in Fig. 1-2; the data are based on Ref. 1. These yields represent the probability of a core hole in the K or L shells being filled by a radiative process, in competition with nonradiative processes. Auger processes are the only nonradiative processes competing with fluorescence for the K shell and

110. The plotted curve for the L shell rep.

average of L_1 , L_2 , and L_3 effective yields.



"X-Ray Data Bqo/det" Section 1.3





Figure 2



MATRIX ELEMENTS IN The Soft and Hard X-Ray Spectroscopies: RESONANT EFFECTS



• Resonant photoemission:

$$I \propto \left| \langle \varphi_{f}(\mathbf{1}) | \hat{e} \bullet \vec{r} | \phi_{i}(\mathbf{1}) \rangle + \sum_{m} \langle \varphi_{f}(\mathbf{1}) \varphi_{1}(\mathbf{2}) | \frac{e^{2}}{r_{12}} | \varphi_{i}(\mathbf{1}) \varphi_{m}(\mathbf{2}) \rangle \langle \varphi_{m}(\mathbf{1}) | \hat{e} \bullet \vec{r} | \varphi_{1}(\mathbf{1}) \rangle \right|^{2}$$

$$\times \delta(h_{V} - (E_{m} - E_{1}))$$
Direct
$$Vacuum - - \bigcap_{\phi_{f}} \overline{Resonant}$$

$$W = \bigcap_{\phi_{i}} \widehat{\varphi_{i}}$$


SrTiO₃ and La_{0.67}Sr_{0.33}MnO₃ band structures and DOS



Resonant Photoemission—La_{0.6}Sr_{0.4}MnO₃, Mn 3d with Mn 2p



K. Horiba et al. | Journal of Magnetism and Magnetic Materials 272–276 (2004) 436–437



MATRIX ELEMENTS IN The Soft and Hard X-Ray Spectroscopies: RESONANT EFFECTS

• X-ray emission:
$$I \propto \left| \hat{\mathbf{e}} \bullet \left\langle \varphi_f(\mathbf{1}) \right| \vec{r} \left| \varphi_i(\mathbf{1}) \right\rangle \right|^2$$



X-ray fluorescence spectroscopy =X-ray emission spectroscopy (XES) and Resonant inelastic xray scattering (RÍXS) u.] g inelastic Intensity x-ray scattering (RIXS) and Resonant elastic x-ray scattering (REXS)



Excitations probed by RIXS



Hancock et al., Phys. Rev. B 80, 092509 (2009)





INTENSITIES IN PHOTOBLECTRON SPEETRA:

· GENERAL: FINAL STATE K (K-SUBSHELL + ALL OTHER DESIG.)

"Basic Concepts of XPS" Section 3.D.

INTENSITIES IN PHOTOBLECTRON SPECTRA:

B GENERAL' FINAL STATE K (K-SUBSHELL + ALL OTHER DESIG)

THT. Kale .< Y tot (N. K) (F F (N)) 2 (DIFOLE APPROX.

. BORN-OMENHEIMER : ET'S FAST, VIBRATIONS SLOW

INT. Kale VIE. V. IVis. VIE (A. K) (E. T. IV. (N))

"Basic Concepts of XPS" Section 3.D.

INTENSITIES IN PHOTOBLECTRON SPECTRA:

GENBRAL' FINAL STATE K (K-SUBSHELL + ALL OTHER DESIG) THT. C C. CY (N. K) ET; I T'(N) /2 (DIFOLE APPROX. BORN-OPPENHEIMER : ET'S FAST, VIBRATIONS SLOW INT. K ((U , K) (I U , K) (I C , K) (I $\frac{\varphi - \varphi}{\varphi_{k+1}} \left(\frac{\varphi}{\varphi_{k+1}} \right) \left(\frac{\varphi$ $INT_{k} \propto |\langle \Psi_{n}^{\dagger}_{0}, v | \Psi_{n}^{i}_{0}, v \rangle|^{2} |\langle \Psi_{2}^{\dagger}(N-1, k) | \Psi_{2}^{\dagger}(N-1, k) \rangle|^{2}$ 2.<4,1 F(4,) SAME SUBSHELL COUPLING + 4 NILMAL de TOTAL L,S→"MONOPOLE"

> "Basic Concepts of XPS" Section 3.D.

INTENSITIES IN PHOTOBLECTRON SPECTRA:

GENERAL: FINAL STATE K (K-SUBSHELL + ALL OTHER DESIG) THT. C O. CHE (N. K) EF. (P'(N)> (2 (DIFOLE APPROX. BORN-OMENHEIMER : ET'S FAST, VIBRATIONS SLOW INT. K ((U , K) (U $k \text{ HOLE} \xrightarrow{\varphi_{k+1}} \begin{array}{c} \varphi_{k-1} \\ \varphi_{k+1} \\ \varphi_{k+1} \end{array} \xrightarrow{\varphi_{k+1}} \begin{array}{c} \varphi_{k-1} \\ \varphi_{k+1} \\ \varphi_{k} \end{array} \xrightarrow{\varphi_{k+1}} \begin{array}{c} \varphi_{k-1} \\ \varphi_{k} \\ \varphi_{k} \end{array} \xrightarrow{\varphi_{k+1}} \begin{array}{c} \varphi_{k-1} \\ \varphi_{k} \\ \varphi_{k} \end{array}$ Ìĉ.< 𝚓 I ₽ (𝐅) > SAME SUBSHELL COUPLING + 4 NERMAL de TOTAL L,S-"MONOPOLE" · SLATER DETS. FOR $\Psi_{e}^{\#} = det(\Psi_{e}^{'}, \Psi_{e}^{'}, \Psi_{e}^{'}, \Psi_{e}^{'})$ $\text{Int.}_{k} \propto |\langle \psi_{v_{10},v}^{\pm}, |\psi_{w_{0},v}^{i} \rangle|^{2} |\langle \psi_{v}^{i}|\psi_{v}\rangle|^{2} |\langle \psi_{v}|\psi_{v}\rangle|^{2} |\langle \psi_{v}|\psi_{v}\rangle|^{2} |\langle \psi_{v}|\psi_{v}\rangle|^{$ $|\langle \Psi_{k-1}|\Psi_{k-1}\rangle|^{2}|\langle \Psi_{k+1}|\Psi_{k+1}\rangle|^{2}\cdots|\langle \Psi_{k}||\Psi_{k+1}\rangle|^{2}$ 12. < 22 17 14 >12 (N-1)e⁻ SHAKE-UP/ SHAKE-OFF \rightarrow **1e- DIPOLE** \rightarrow d σ /d Ω "MONOPOLE"

"Basic Concepts of XPS" Section 3.D.

INTENSITIES IN PHOTOBLECTRON SPECTRA!

GENERAL: FINAL STATE K (K-SUBSHELL + ALL OTHER DESIG) THT. C O. CHE (N. K) EF. (P'(N)> (2 (DIFOLE APPROX. BORN-OPPENHEIMER : ET'S FAST, VIBRATIONS SLOW INT. K ((U , K) (U $\begin{array}{c} \varphi_{i} \rightarrow \varphi_{i}^{\prime} \\ \varphi_{k+1} \rightarrow \varphi_{k+1}^{\prime} \end{array} \qquad (Soow) \\ \varphi_{k+1} \rightarrow \varphi_{k} \\ \varphi_{k} \rightarrow \varphi_{k} \end{array}$ le.<(4,1 F(4,)) SAME SUBSHELL COUPLING + 4 NICCHAR 46 TOTAL L,S→"MONOPOLE" · SLATER DETS. FOR Ye = dat (4'4' ... 4' 4' ... 4') $\text{Int.}_{k} \propto |\langle \psi_{v_{10},v}^{\pm}, |\psi_{w_{0},v}^{i} \rangle|^{2} |\langle \psi_{v}^{i}|\psi_{v}\rangle|^{2} |\langle \psi_{v}|\psi_{v}\rangle|^{2} |\langle \psi_{v}|\psi_{v}\rangle|^{2} |\langle \psi_{v}|\psi_{v}\rangle|^{$ $|\langle \Psi_{k-1}|\Psi_{k-1}\rangle|^{2}|\langle \Psi_{k+1}|\Psi_{k+1}\rangle|^{2}\cdots|\langle \Psi_{k}||\Psi_{k}\rangle|^{2}$ 12. < 92 17 19, >12 (N-1)e⁻ SHAKE-UP/ "Basic Concepts of XPS" SHAKE-OFFightarrow**1e- DIPOLE** \rightarrow d σ /d Ω Section 3.D. "MONOPOLE"

• PLUS DIFFRACTION EFFECTS IN 4 ESCAPE : PD or EXAFS

The vacuum ultraviolet, soft x-ray, hard x-ray measurements:

The spectroscopies: free

hν

hν

PHOTOELECTRON SPECTROSCOPY= PHOTOEMISSION – PS, PES, UPS, XPS + DIFFRACTION-XPD, PhD + HOLOGRAPHY-PH + MICROSCOPY-PEEM



X-RAY ABSORPTION SPECTROSCOPY- XAS *NEAR-EDGE – NEXAFS, XANES + X-RAY MAGNETIC CIRCULAR/LINEAR DICHROISM- XMCD, XMLD #EXTENDED- EXAFS, XAFS



X-RAY EMISSION (FLUORESCENCE) SPECTROSCOPY

+ AUGER ELECRON SPECTROSCOPY (Always accompanies photoelectron emission) 121





Scanned-energy photoelectron diffraction—an alternative approach (Shirley, Woodruff/ Bradshaw, Lapeyre, Chiang et al.)



Photoelectron diffraction: Simple singlescattering theory for ssubshell emission

Various papers by Shirley et al., Woodruff, Bradshaw et. al.



COMPARISON OF SCANNED-ENERGY PD TO EXTENDED X-RAY ABSORPTION FINE STRUCTURE



Figure 1.1. The x-ray absorption coefficient for the K-edge of copper metal.

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+ COPPER
15 DINDING
ENERGY
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Also scanned-energy, but integrates over all electron emission directions



Theory of Extended X-Ray Absorption Fine Structure









Theory of Extended X-Ray Absorption Fine Structure



COMPARISON OF SCANNED-ENERGY PD TO EXTENDED X-RAY ABSORPTION FINE STRUCTURE



Figure 1.1. The x-ray absorption coefficient for the K-edge of copper metal.

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    Also scanned-energy,
    COPPER but integrates over all electron emission
    ENERGY directions
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The quantum mechanics of covalent bonding in molecules: H_2^+ with one electron



IGURE 8.4 (a) Potential energy of an electron in the electric field of two nearby protons. The total energy of a ground-state electron in the hydrogen atom is indicated. (b) Two nearby protons correspond quantum-mechanically to a pair of boxes separated by a barrier.

NIL MOLECULAR ORBITAL DRAWINGS

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 $10_{0} \epsilon = -0.5944 \text{ a.u.} = -16.16 \text{ eV}$ 136 (Compare – 13.61 for H atom 1s)





THE ELECTRONS IN HF (OR HCI): ionic molecules

1 e⁻



E = -25.6 a.u.

1σ

PHOTOELECTRON EMISSION-

BASIC MATRIX ELEMENTS + SELECTION RULES:



PHOTOELECTRON EMISSION-

BASIC MATRIX ELEMENTS + SELECTION RULES:







NEARLY-FREE ELECTRONS IN A WEAK PERIODIC POTENTIAL-1 DIM.



NEARLY-FREE ELECTRONS IN A WEAK PERIODIC POTENTIAL-1 DIM.



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Figure 28 Brillouin zones of the face-centered cubic lattice. The cells are in reciprocal space, and the reciprocal lattice is body-centered, as drawn.

- STACKING OF FCC BRILLOUIN ZONES -



Electronic bands and density of states for a <u>semiconductor</u>-Germanium— 1s²2s² 2p⁶3s² 3p⁶3d¹⁰4s²4p²



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Fig. 7.12. Bandstructure E(k) for copper along directions of high crystal symmetry (*right*). The experimental data were measured by various authors and were presented collectively by Courths and Hüfner [7.4]. The full lines showing the calculated energy bands and the density of states (*left*) are from [7.5]. The experimental data agree very well, not only among themselves, but also with the calculation



And for the real d orbitals in fcc:



Ligand (e.g. O)

X

Χ

Copper densities of states-total and by orbital type:









Hathaway et al., Phys. Rev. B 31, 7603 ('85)

Fe: ANGLE AND SPIN-RESOLVED SPECTRA AT Γ POINT



And for the real d orbitals in octahedral:





E.g.—Crystal field in Mn³⁺ & Mn²⁺ with negative octahedral ligands



SrTiO₃ and La_{0.67}Sr_{0.33}MnO₃ band structures and DOS





SURFACE ELECTRONIC STATES



STRUNGLY LOCALIZED NEAR SURFACE

BLOCH FUNCTION IN X+Y, BUT DECAYING IN Z:





ONLY EXIST WHEN NO BULK STATE
EXISTS AT SAME En = Lxî+Lyĵ; OTHERWISE
MIXING OCCURS + NOT SURFACE-LOCALIBED

Surface states on Cu(111)

Fig. 4.21. Experimental dispersion of Cu(111) surface states plotted with a projection of the bulk bands: (a) Shockley state near the zone center (Kevan, 1983); (b) Tamm state near the zone boundary (Heimann, Hermanson, Miosga and Neddermeyer, 1979). Compare with Fig. 4.17.



CONSERVATION LAWS IN VALENCE-BAND PHOTOBLECTRON SPECTROSCOPY:



Fig. 4.20. Photoemission energy distribution curves from Cu(111) at different collection angles. Equation (4.32) has been used to express the electron kinetic energy in terms of the binding energy of the electron state (Kevan, 1983).



Graphene- A very special 2D case



Bostwick et al., Nature Physics 3, 36 - 46(2007)

REXS, XRD, XAS, Some basic measurements: **NEXAFS/XANES**, **XES, RIXS XFH** Standing $\lambda_{sw} = \lambda_x / (2 \sin \theta_x)$ EXAFS, XRO XFH⁻¹, RXFH Wave: LP, RCP, LCP, UP hv' PS, PD, PH <h CD, MCD, SP, I λ_{SW} exp(-μ_xL_x)= $exp(-L_{o}/\Lambda_{o})$ $exp(-L_x/\Lambda_x)$ $n = 1 - \delta - i\beta$ (Kramers-Kronig) fx X-ray Fluorescence Holography (XFH, XFH¹), Resonant XFH (RXFH) ١X \approx 1-(r₀ $\lambda_{x}^{2}/2\pi)\Sigma n_{i}f_{xi}(0)$ X-ray Emission Spectroscopy (XES), $\mu_{\rm X} = 4\pi\beta/\lambda_{\rm X}$ **Resonant Inelastic X-ray Scattering (RIXS)** Multi-atom resonant **Resonant Elastic X-ray Scattering (REXS)** photoemission (MARPE) $\theta_{x}^{R} = \theta_{x}^{I}$ Photoelectron X-Ray Diffraction (XRD) Spectroscopy (PS). $\theta_{CRIT}^{I} = (2\delta)^{1/2}$ X-ray Absorption Spectroscopy (XAS) raction (PD X-Ray Optical measurements (XRO) Holography (PH) Near-Edge X-Ray Absorption Fine-Structure + Circular Dichroism (CD Magnetic CD (MCD), (NEXAFS)/ X-Ray Absorption Near-Edge $R_{x} + T_{x} = 1$ Spin Polarization (SP) Structure (XANES); Extended X-Ray 65 Resonant Photoemission Absorption Fine Structure (EXAFS) T_X